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# Semi-empirical models of X-ray AGN in galaxy clusters:

## The role of environment on AGN triggering

Final conference Bid4BEST - 06/02/2024

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Antonis Georgakakis  
Francesco Shankar

# AGN

## NGC1365



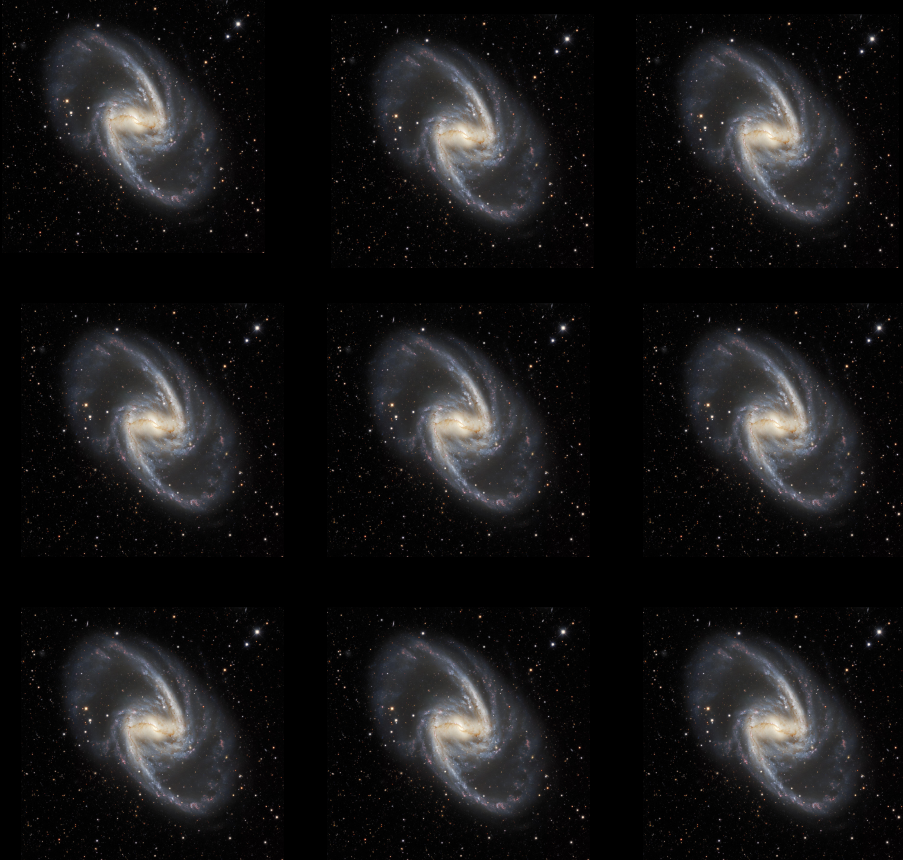
Credits: Dark Energy Survey/DOE/FNAL/DECam/CTIO/NOIRLab/NSF/AURA  
Image processing: Travis Rector (University of Alaska Anchorage/NSF's NOIRLab),  
Jen Miller (Gemini Observatory/NSF's NOIRLab), Mahdi Zamani & Davide de  
Martin (NSF's NOIRLab)

## UGC 6093



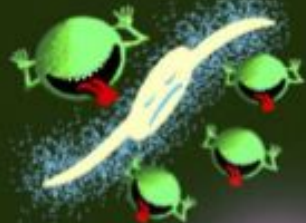
Credits: ESA/Hubble & NASA

Non-AGN more frequent than AGN



Galaxies can go through...  
life is tough!!!

harassment



tidal truncation



ram-pressure stripping



thermal evaporation



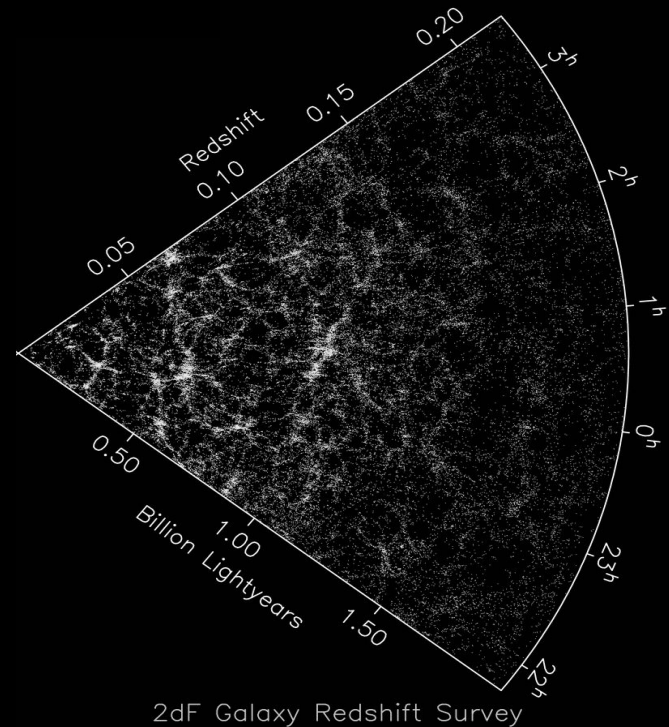
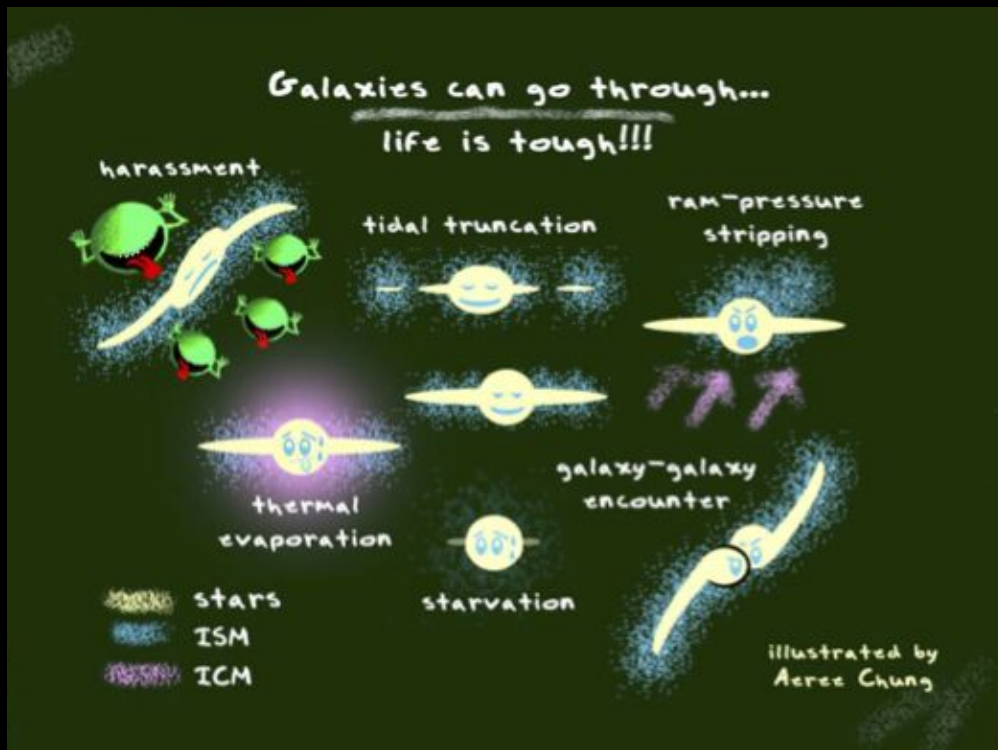
starvation

galaxy-galaxy encounter

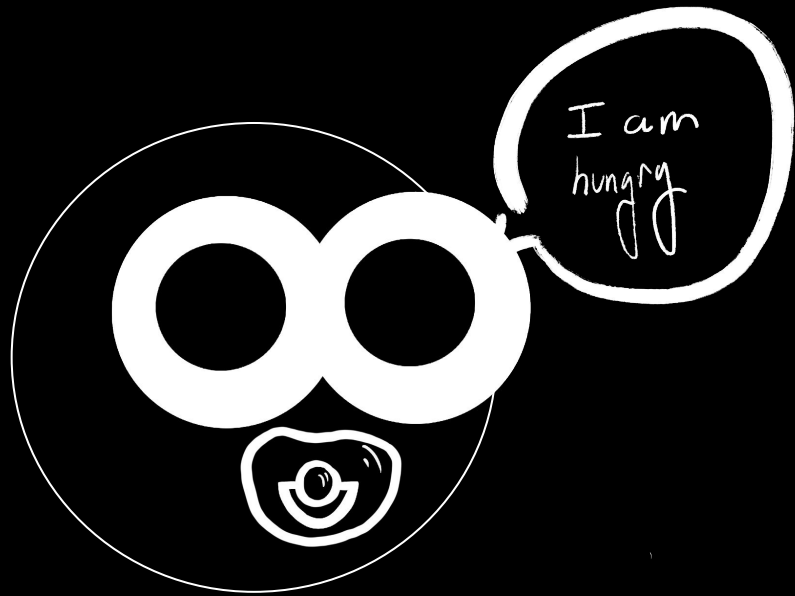


 stars  
 ISM  
 ICM

illustrated by  
Aerec Chung



Yes, life it is tough...but it also depends on the neighbourhood!

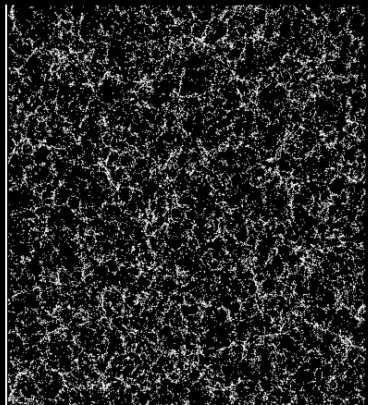


I am  
hungry

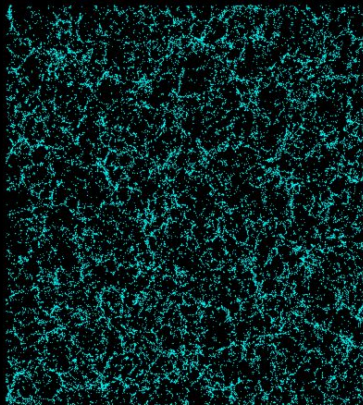
Semi-empirical model

Dark Matter

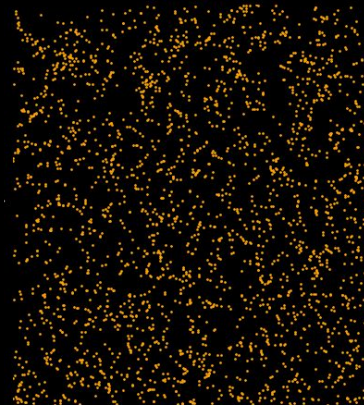
1 000 Mpc/h



Stellar Mass

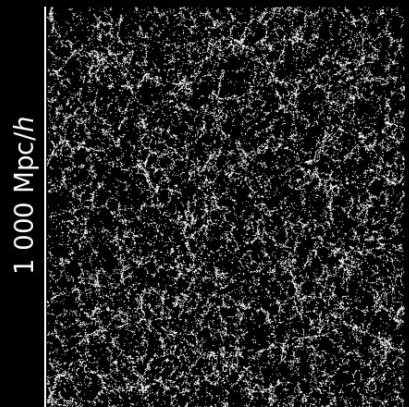


AGN

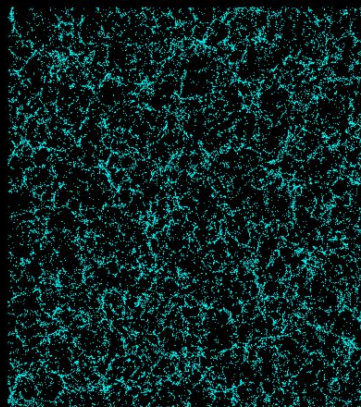


# Semi-empirical model

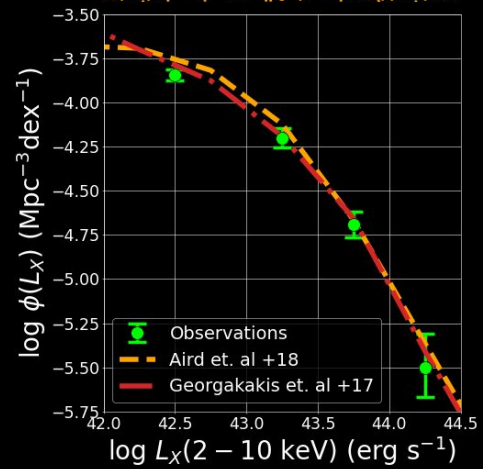
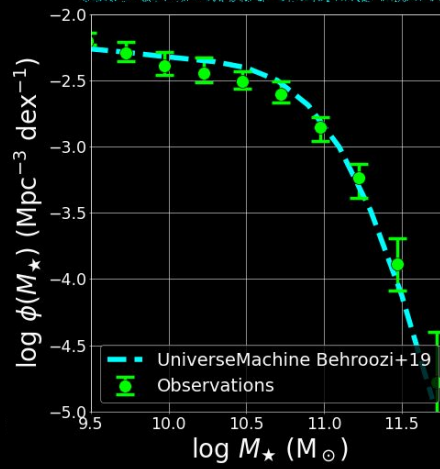
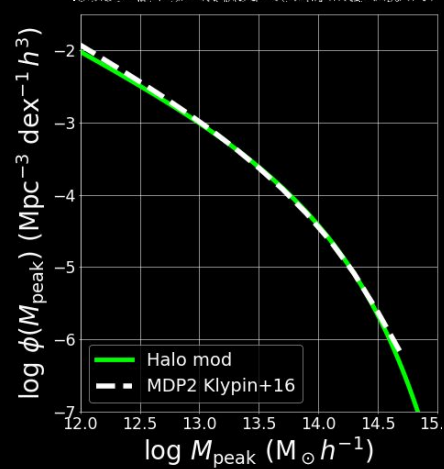
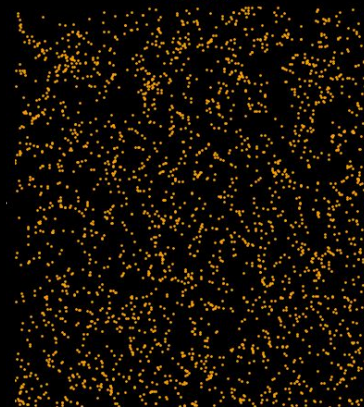
## Dark Matter



## Stellar Mass



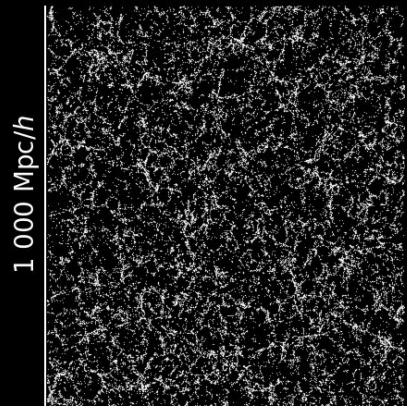
## AGN



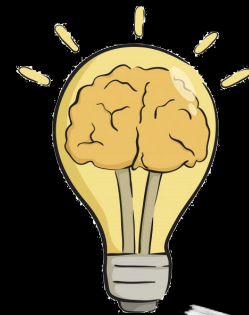
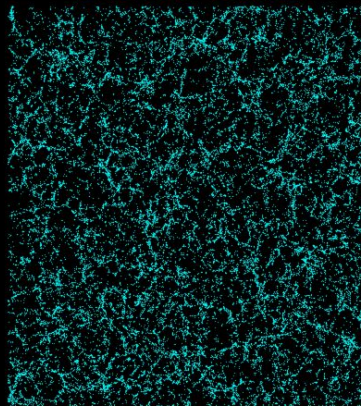


# Semi-empirical model

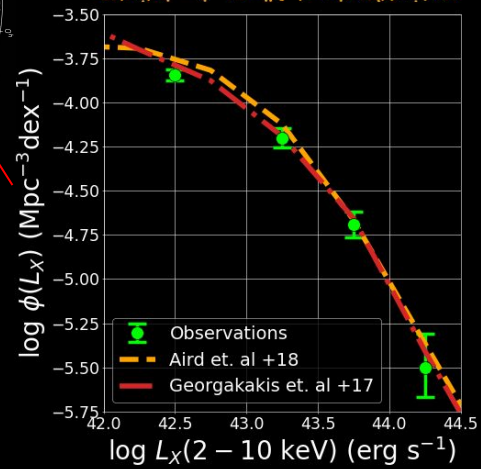
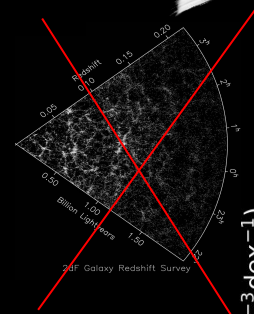
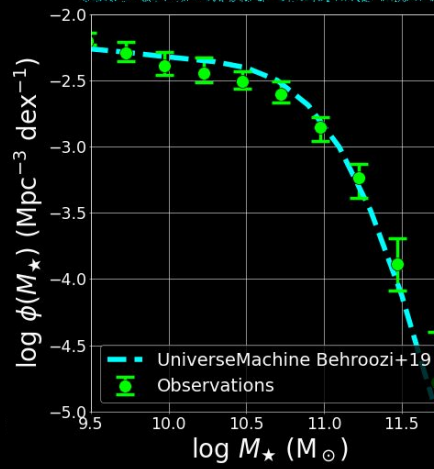
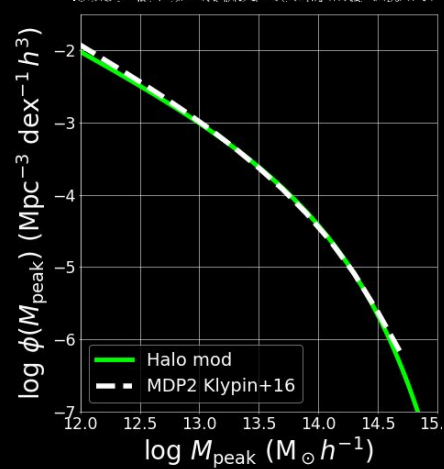
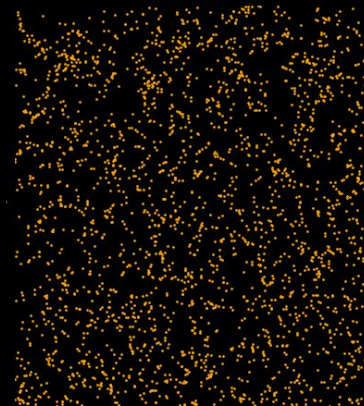
## Dark Matter



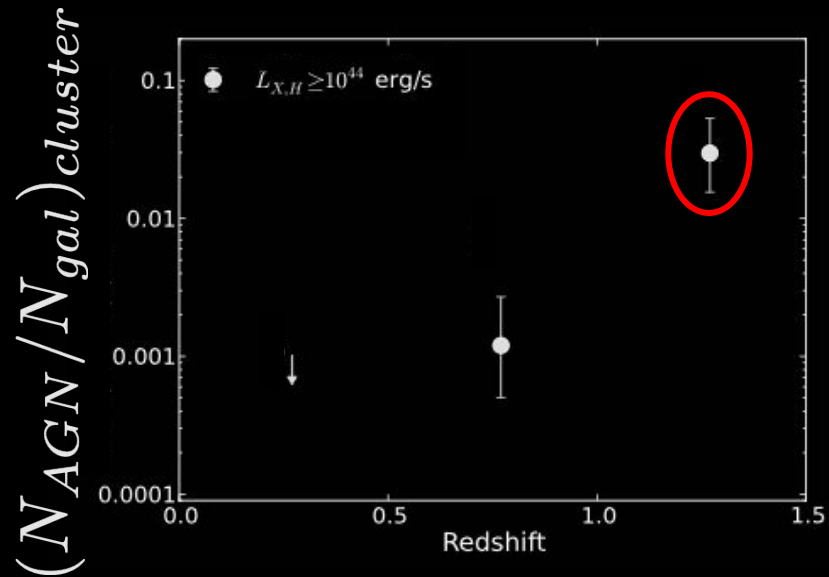
## Stellar Mass



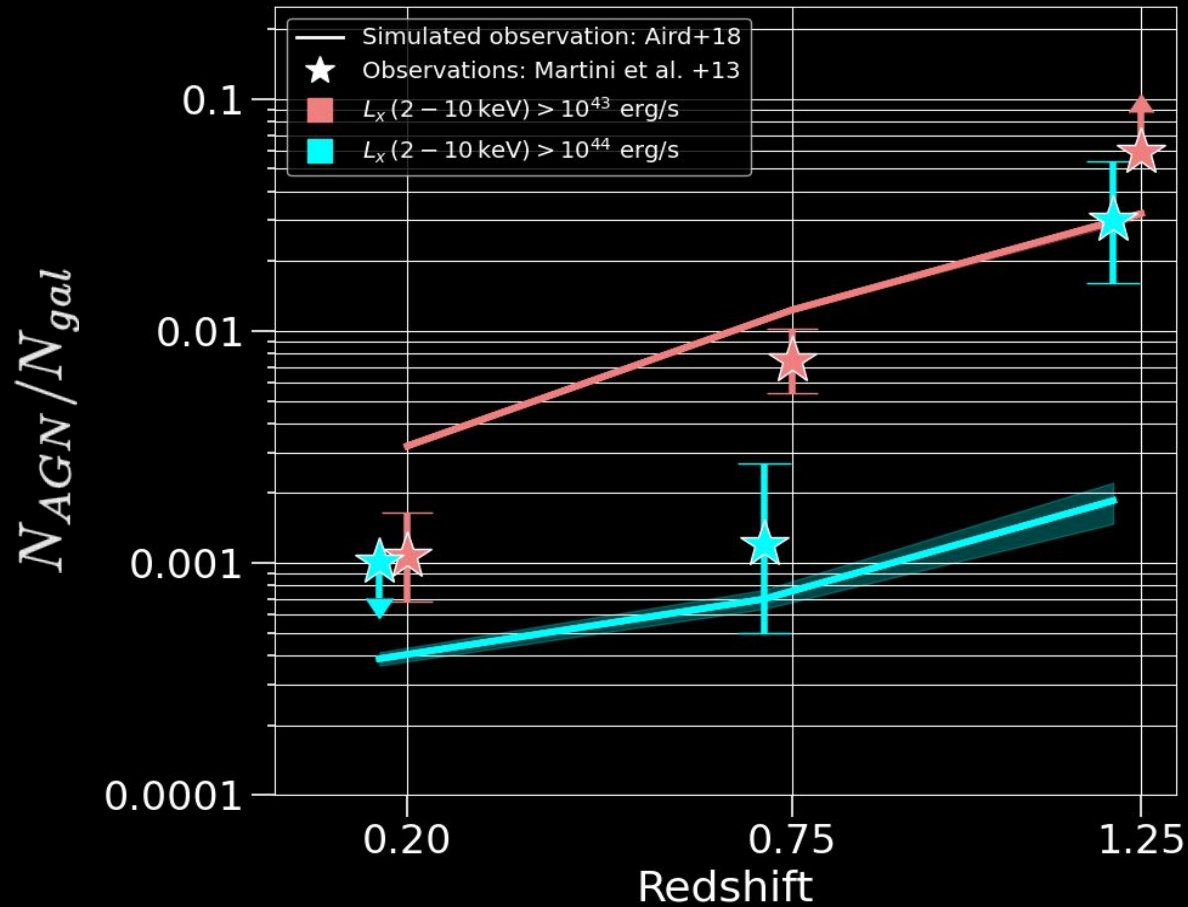
## AGN



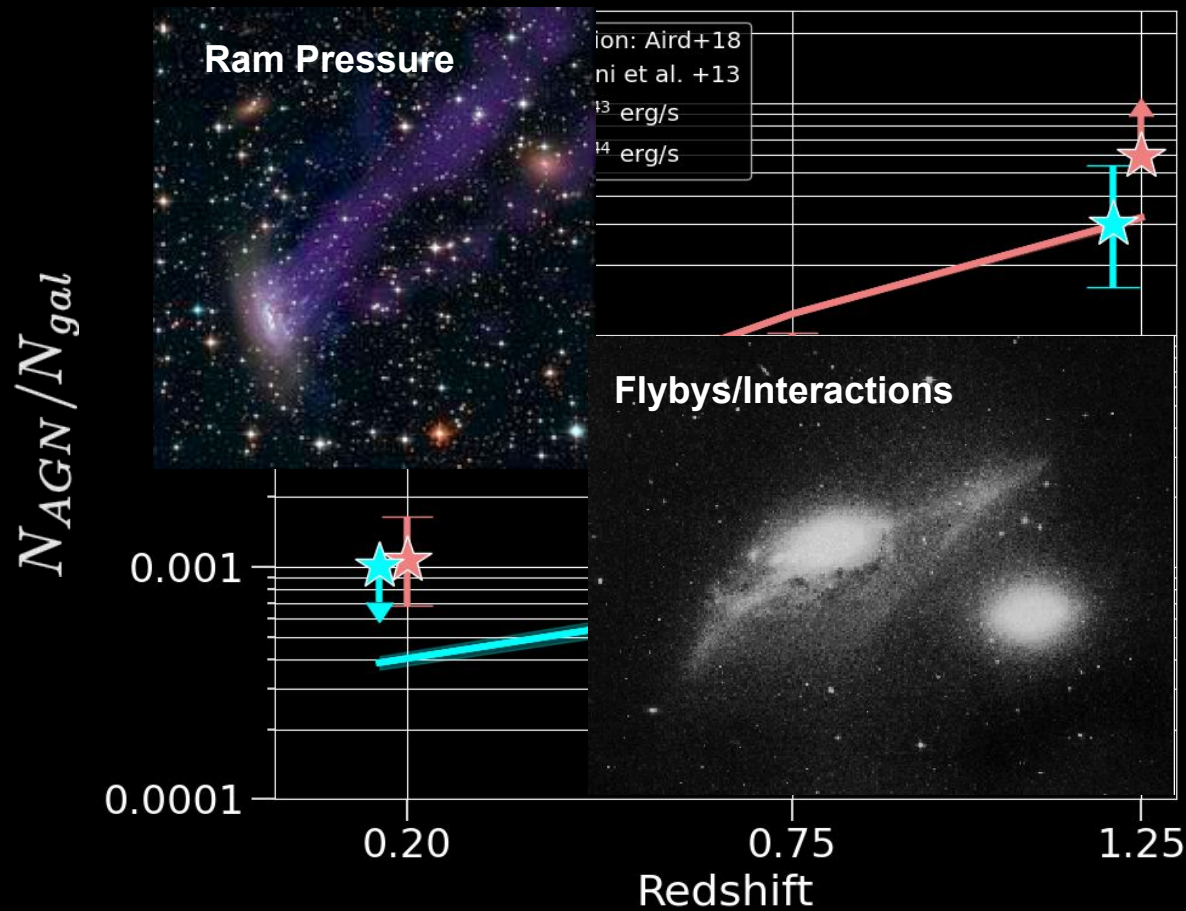
# AGN fractions Martini+13



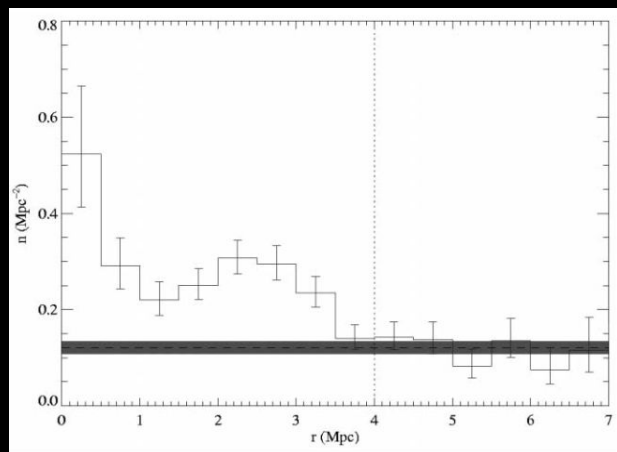
# AGN fractions Muñoz Rodríguez+23



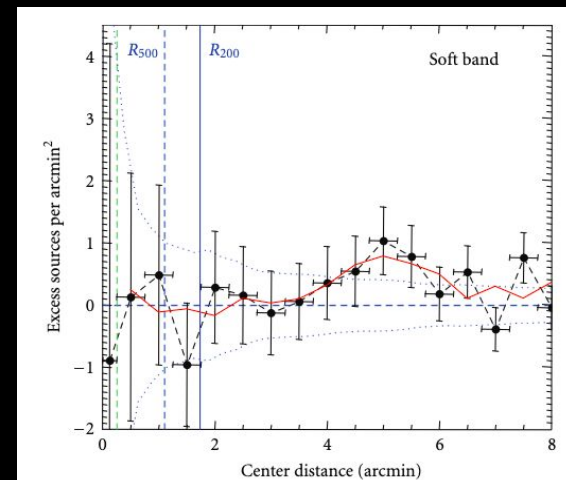
# AGN fractions Muñoz Rodríguez+23



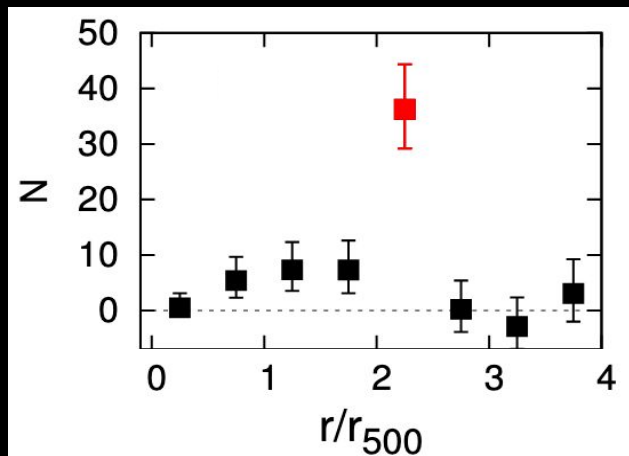
# (Projected) Radial distribution of AGN in clusters



Ruderman & Ebeling (2005)



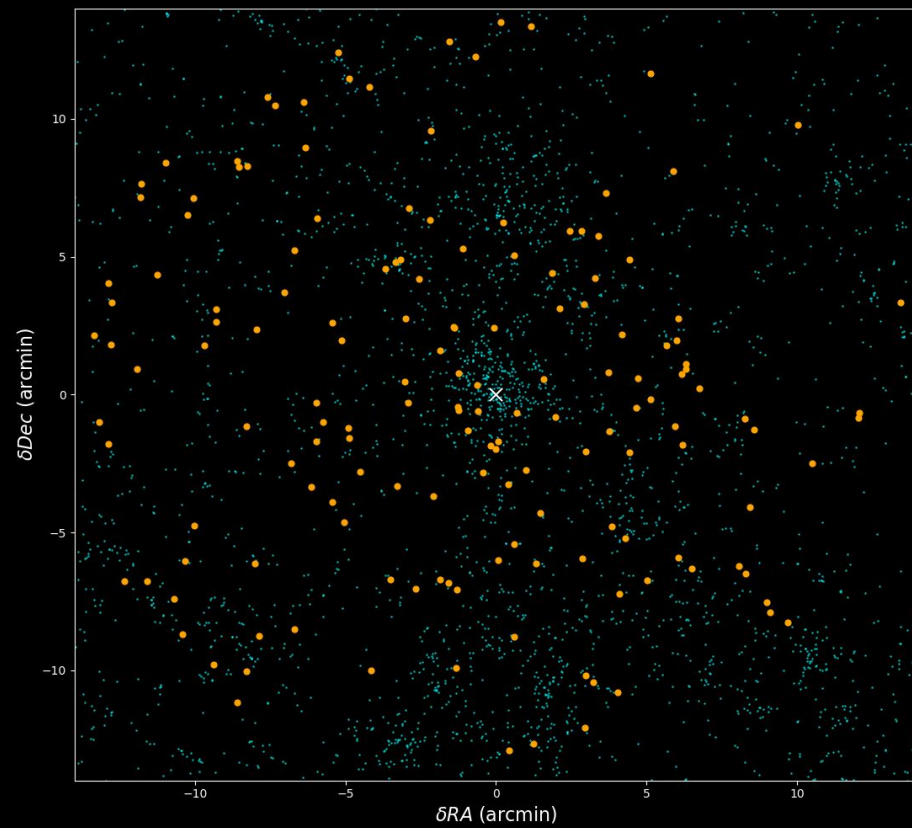
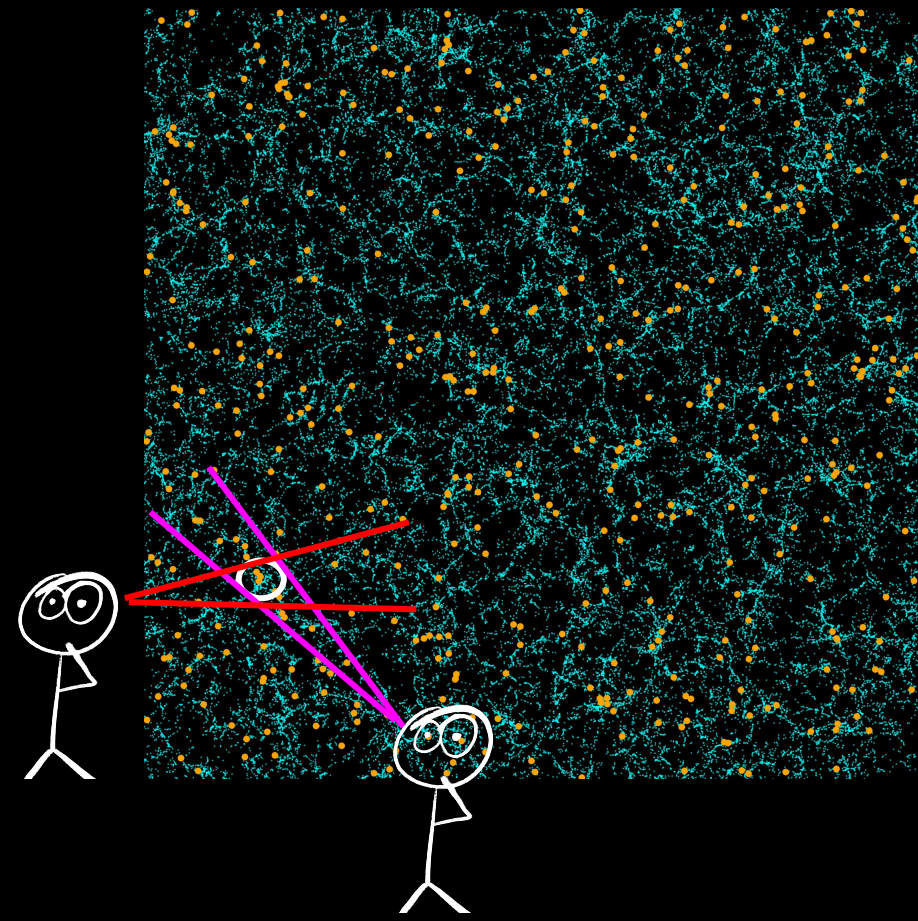
Fassbender, Suhada & A. Nastasi (2012)



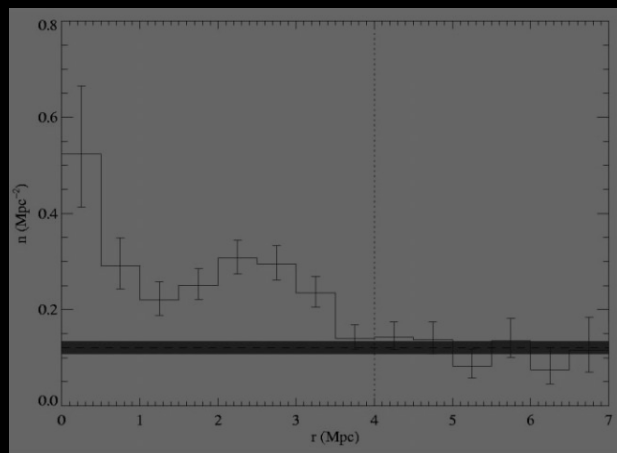
Koulouridis & Bartalucci (2019)

Projection effects?

Galaxies, AGN

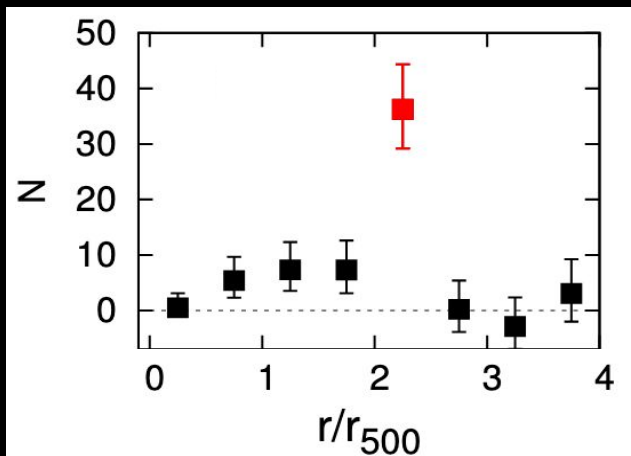


# (Projected) Radial distribution of AGN in clusters

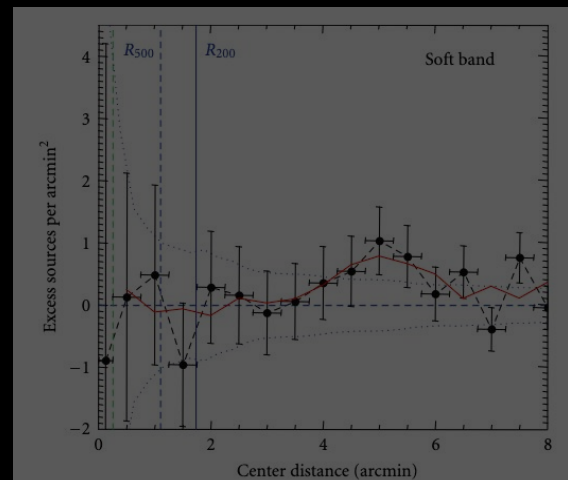


Ruderman & Ebeling (2005)

Mean of 5 clusters

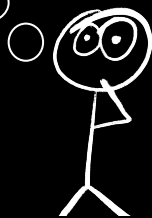
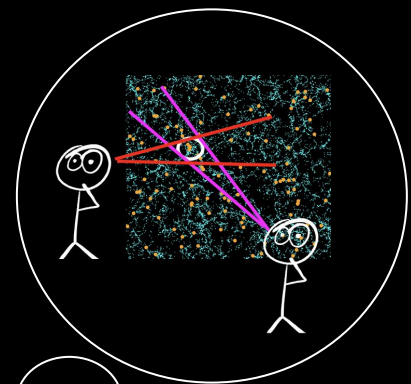
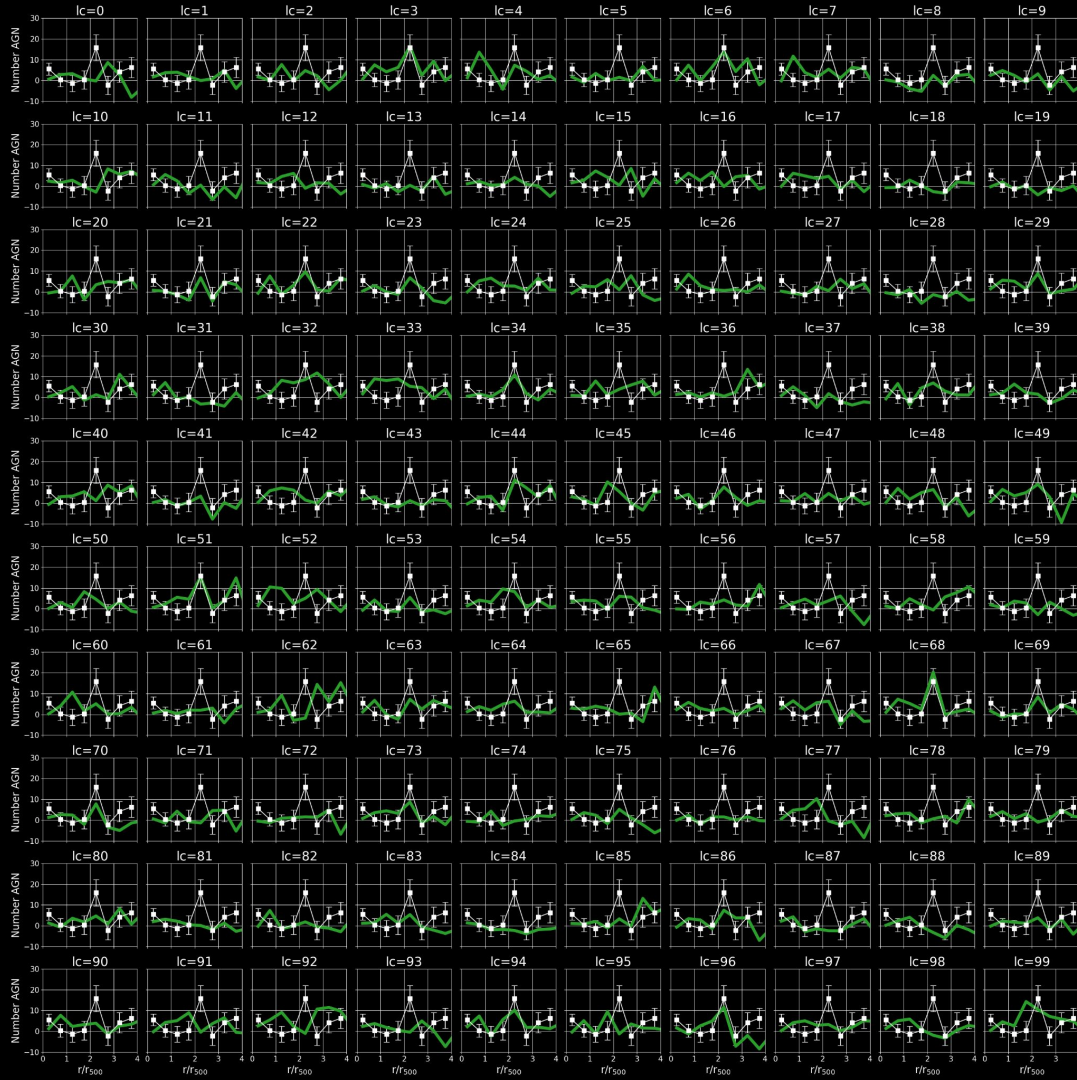


Koulouridis & Bartalucci (2019)



Fassbender, Suhada & A. Nastasi (2012)

# Radial distribution of AGN in clusters

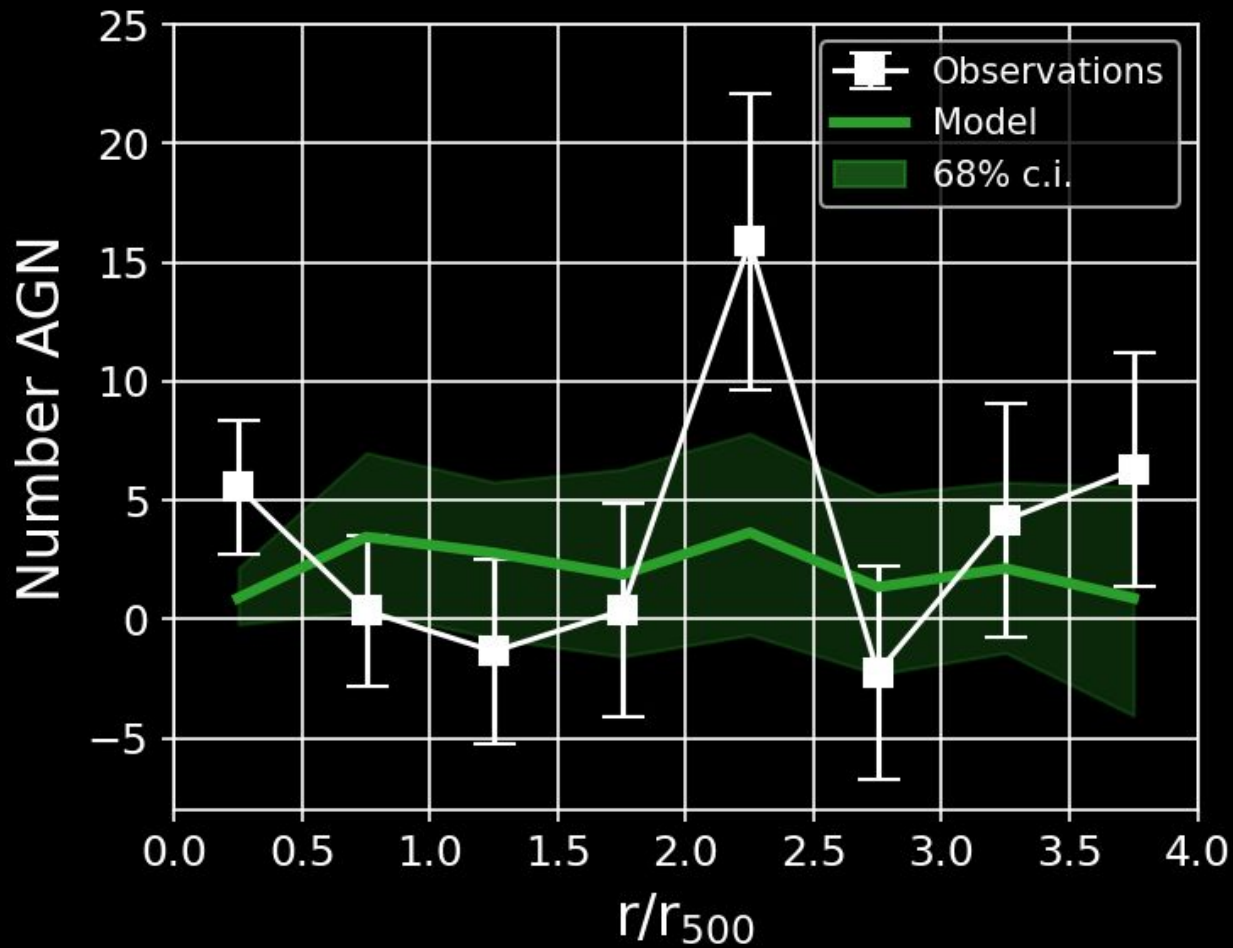


x100

In prep.

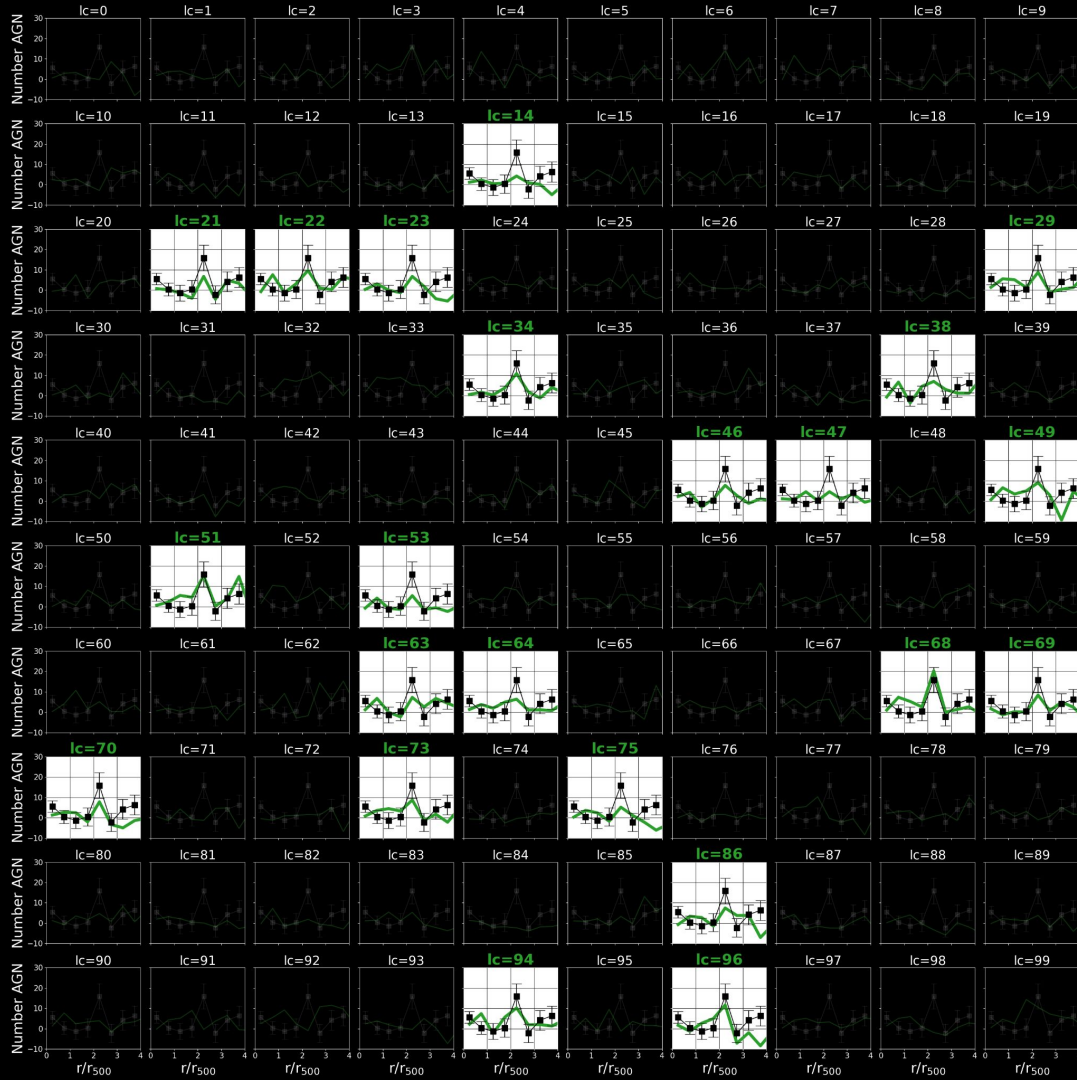


# Radial distribution of AGN in clusters



In prep.

# Radial distribution of AGN in clusters

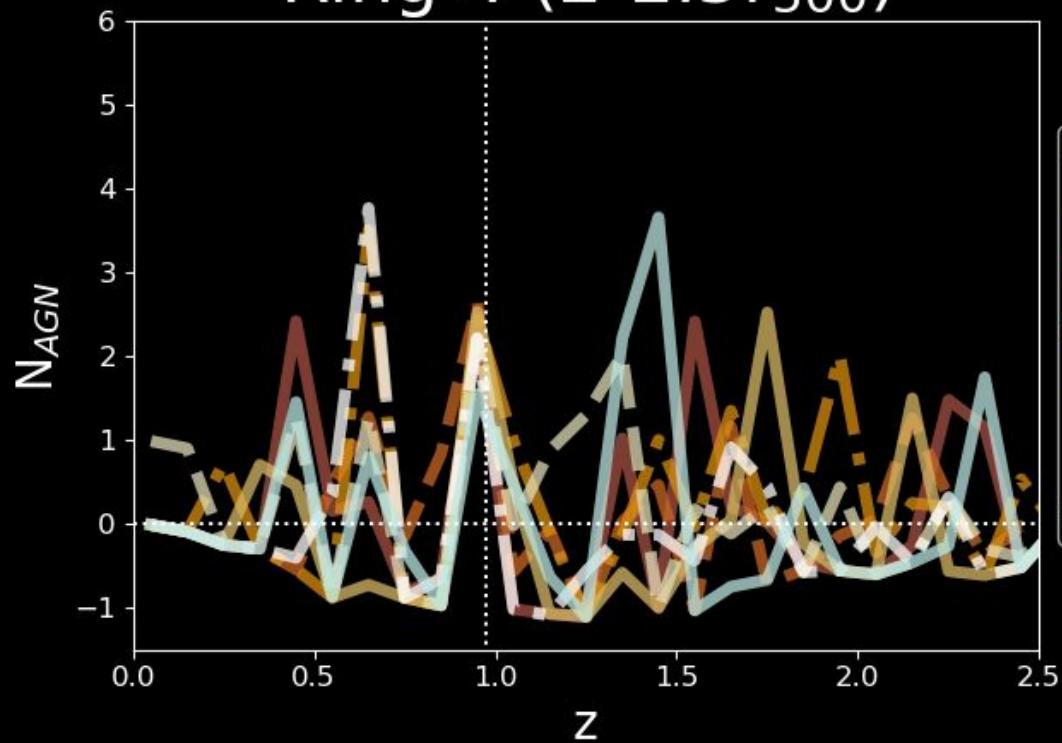


**~ 10-20 % of  
the observers  
show  
an overdensity**

**In prep.**

Radial distribution of  
AGN in clusters

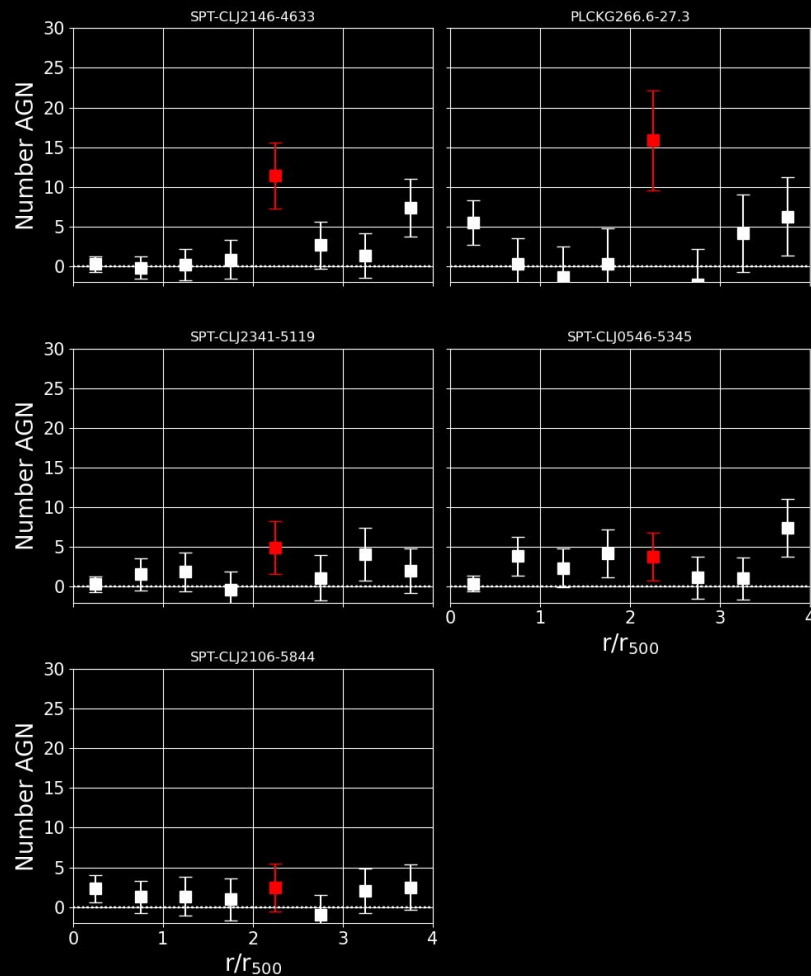
# Ring 4 ( $2-2.5r_{500}$ )



**~ 10-20 % of  
the observers  
show  
an overdensity**

**In prep.**

# Radial distribution of AGN in clusters



**$\sim 50 \pm 20$  % of the clusters observed show a clear overdensity**



## Observations

- Increase sample of cluster and study their radial distributions (on going)
- Optical follow-up (photo. and spect.) of clusters with overdensities (on going)
- Archival data cluster cluster catalogues + spectroscopy (e.g., Abell+SDSS, GOGREEN)
- Explore other environments (e.g., voids)



## Modelling

- Explore modifications of the SEM (including some environment dependence e.g., close pairs)
- Compare these results with other approaches (e.g., hydrodynamical simulations, SAMs)

# THANK YOU!

## PhD: expectation



## PhD: reality





*That's all Folks!*