

# HIGH-Z QUASARS IN MULTI-WAVELENGTH SURVEYS: CONNECTING THE PC TO KPC SCALES

Manda Banerji

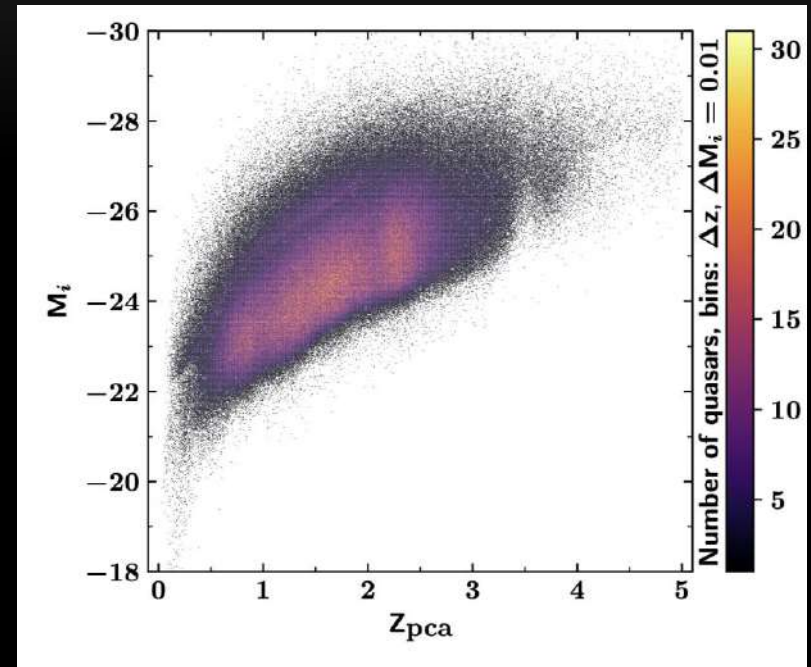
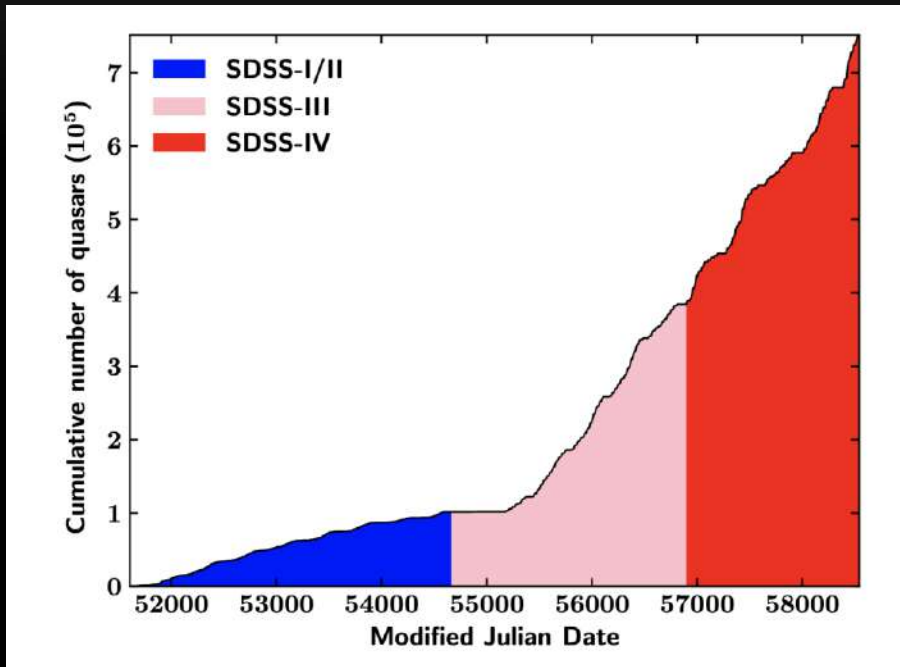
(University of Southampton)

**Matt Stepney (Southampton), Shenli Tang (Southampton), Stephen Molyneux (Southampton)**, Matthew Temple (UDP Chile), Amy Rankine (U. Edinburgh), Paul Hewett (IoA, Cambridge), Gordon Richards (Drexel), James Matthews (Oxford)

# OUTLINE

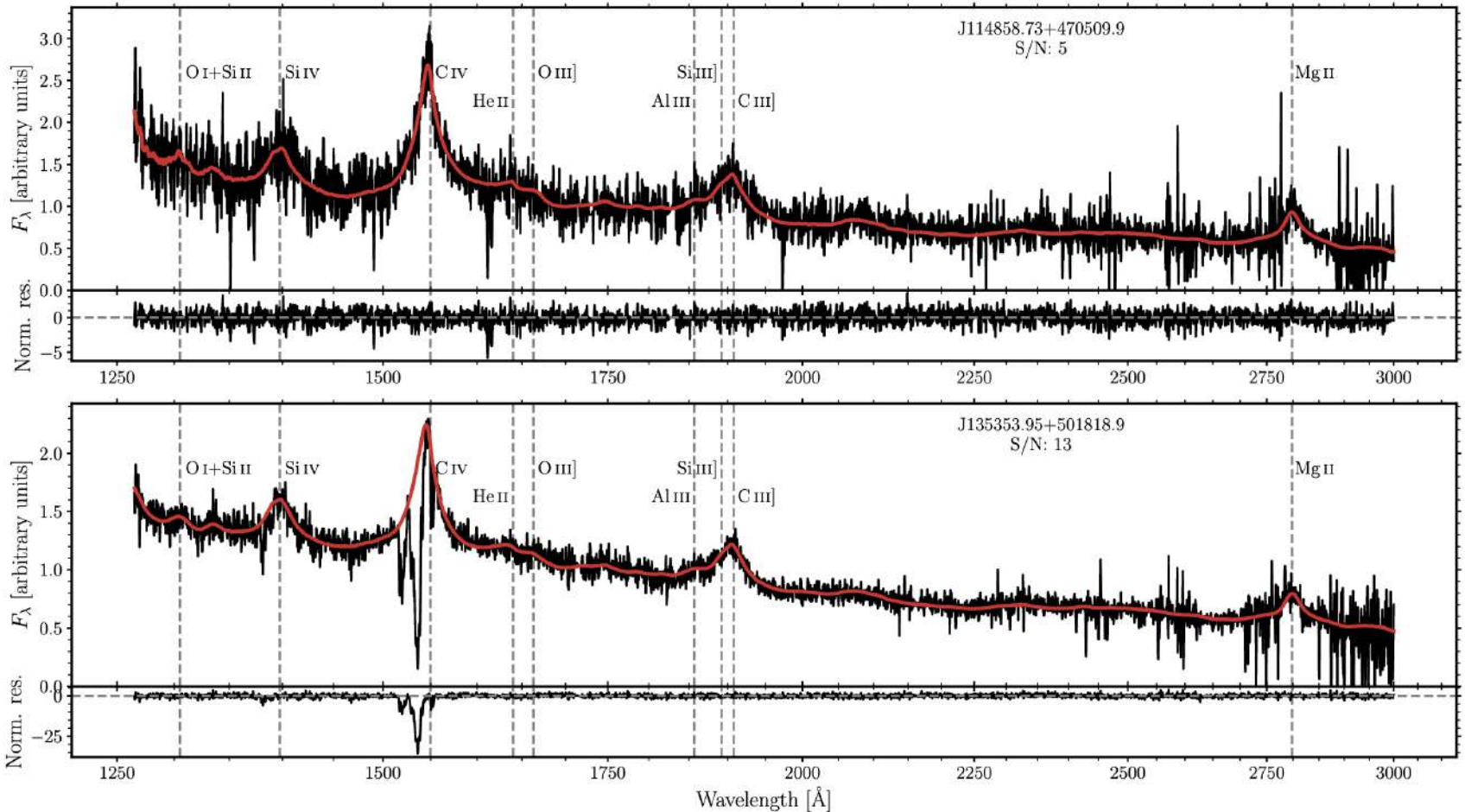
- **High-z quasar demographics from large spectroscopic surveys**
    - Measuring accretion and outflow properties
    - Different tracers of outflows - broad emission lines, broad absorption lines, narrow emission lines
    - Evolution of outflow properties with luminosity, black hole mass, redshift
    - Molecular gas properties
  - **Dusty and red quasars**
    - Differences in outflow properties between red and blue quasars?
    - Connecting black-hole and host galaxy properties via imaging, spectroscopy and molecular gas
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# THE SLOAN DIGITAL SKY SURVEY (SDSS)



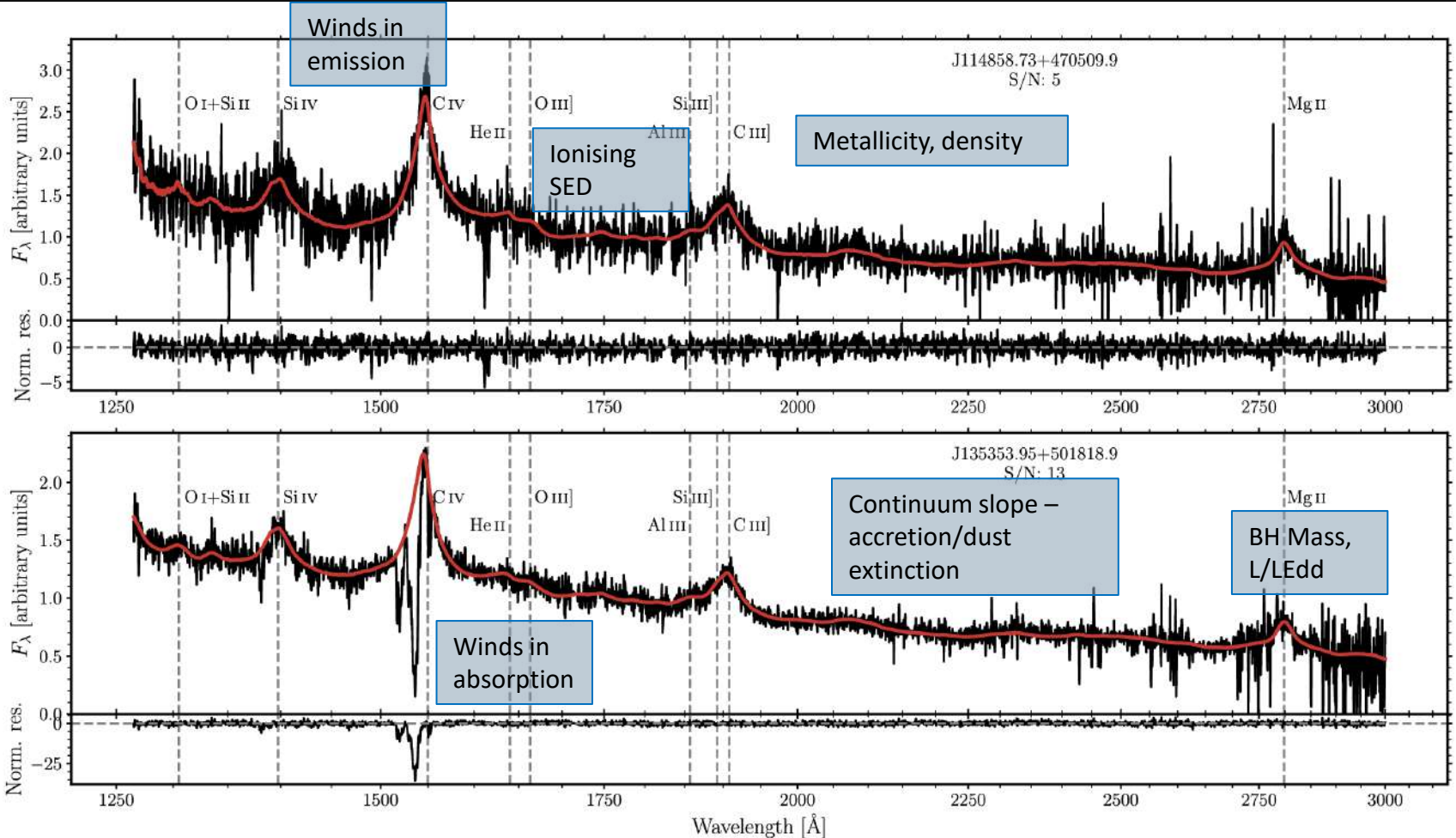
Data Release 16 contains spectra for 750k quasars covering an unprecedented range in luminosity and redshift (Lyke+20) – statistical studies of the high-redshift quasar population now possible

# REST-UV SPECTRA OF HIGH-Z QUASARS



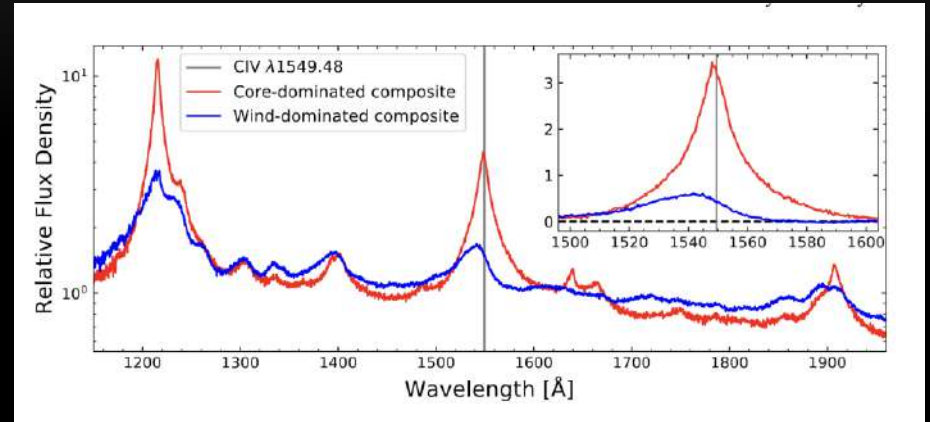
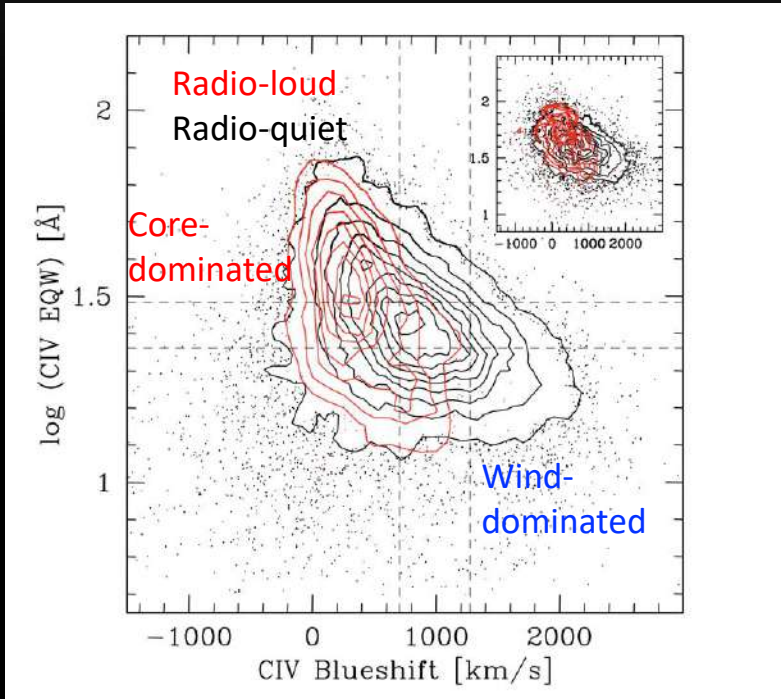
Richards+11; Shen+11; Coatman, Hewett, MB+16,17; Rankine (incl. MB)+20,21; Rakshit+20; Rivera+21,22; Temple, MB+21; Wu&Shen 22; Temple (incl. MB)+23; Stepney, MB+23;

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# THE CIV EMISSION LINE SPACE

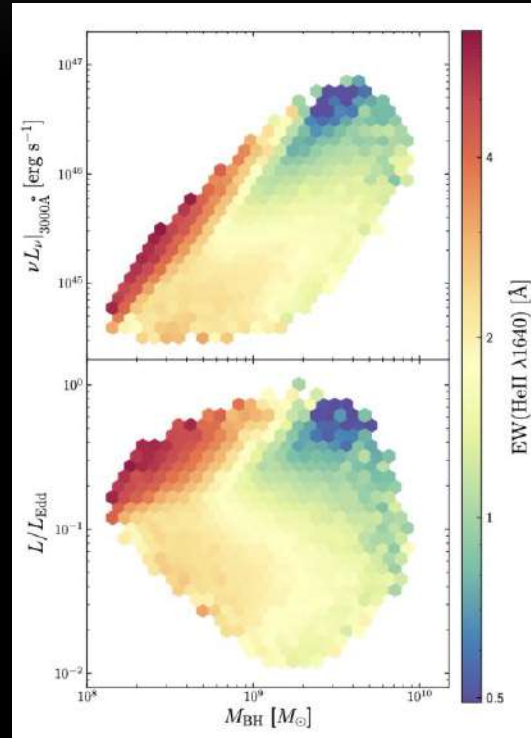
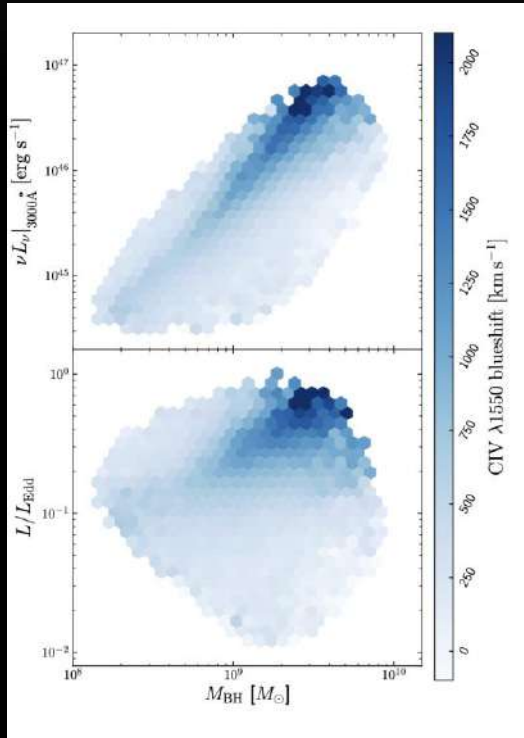


Quasars show systematic trends in the CIV emission line space of blueshift versus EW suggesting this space is tracing fundamental differences in quasar SED and black-hole/accretion properties

Richards+11, Rankine+20,21, Temple+21a,b, Temple+23

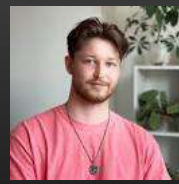
# SUB-PC SCALE WINDS DEPEND ON ACCRETION PROPERTIES

Temple, Matthews..MB..+23



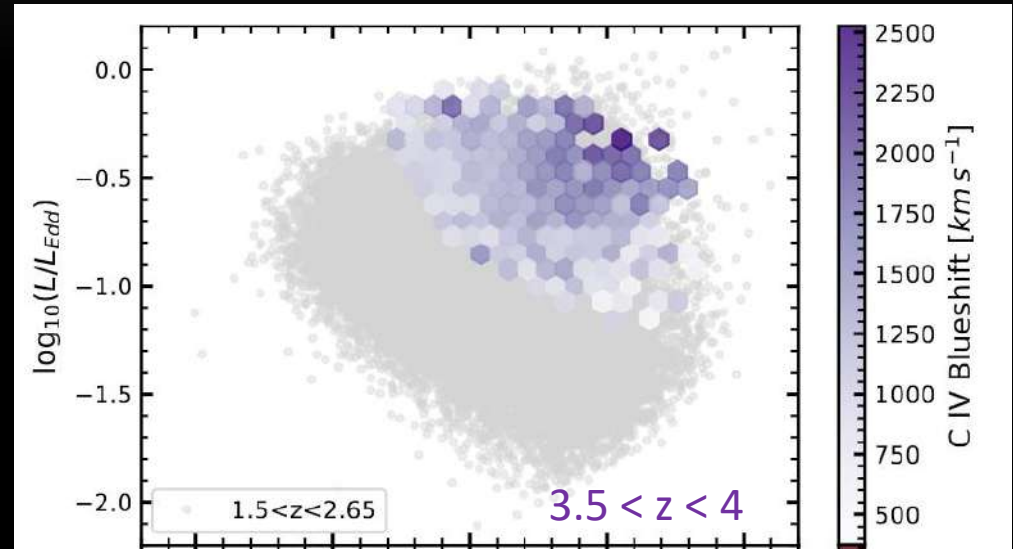
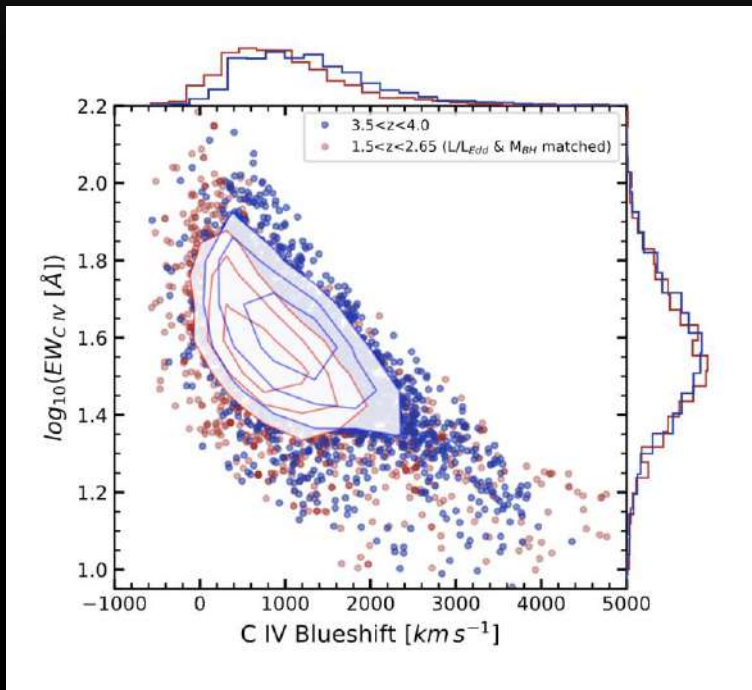
High blueshifts (strong radiatively-driven winds) observed at high black-hole masses and above a critical Eddington ratio of  $>0.25$  consistent with theoretical predictions (e.g. Guistini & Proga 19)

For low MBH and high accretion rates we see increased HeII emission – harder ionizing SED and more line emission due to a failed wind?



# EVOLUTION IN EMISSION LINE PROPERTIES WITH REDSHIFT

Stepney, MB+23



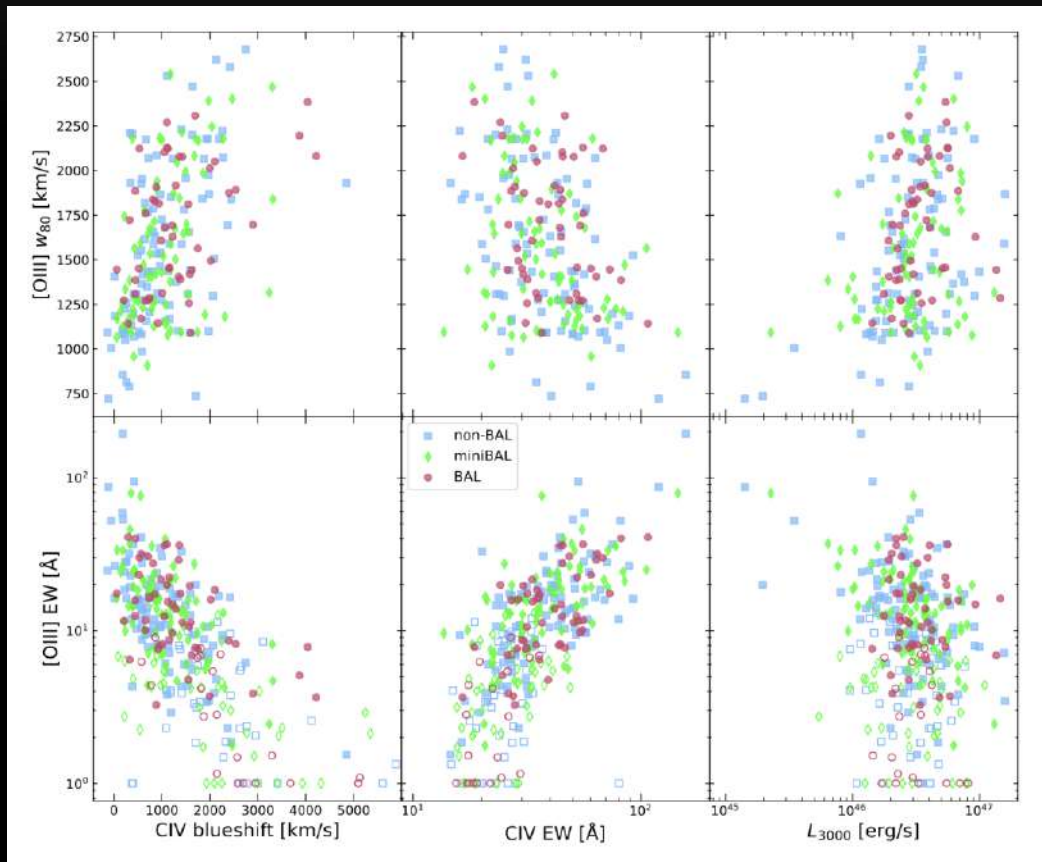
At  $z \sim 4$  (where statistical studies are possible) – no evidence for any redshift evolution in the UV line properties of quasars when we match in fundamental properties like black-hole mass and accretion rate.

4MOST AGN Survey (GTO) will obtain spectra for tens of thousands of  $z > 4$  quasars starting in 2025 – also DESI, WEAVE



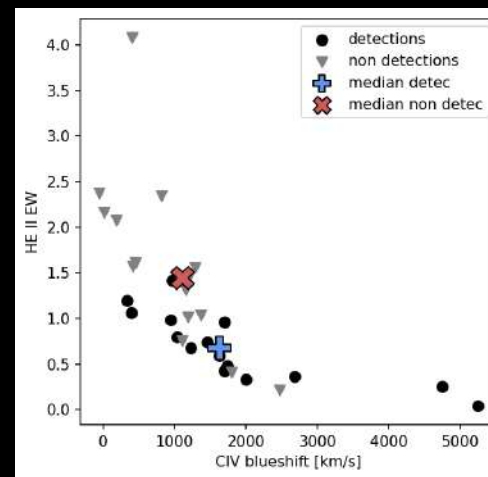
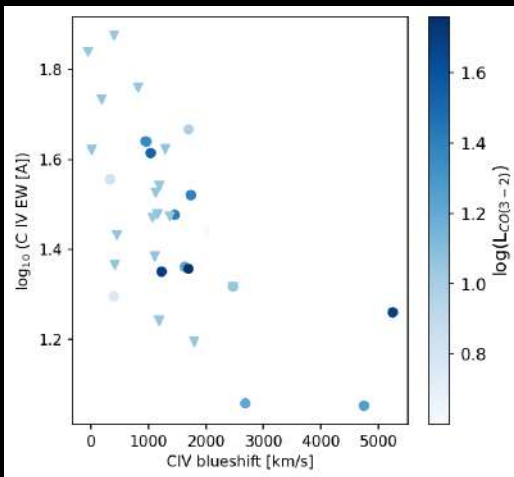
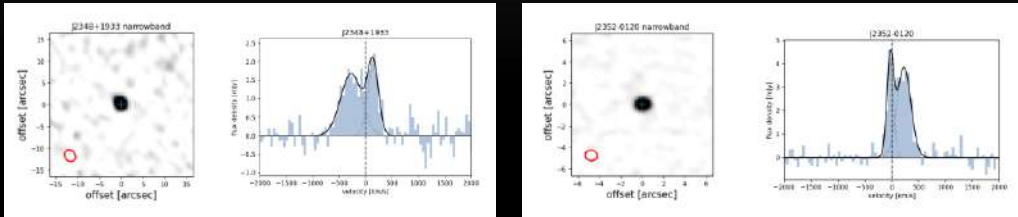
# KILOPARSEC SCALE OUTFLOWS VIA [OIII]

Temple, Rankine, MB+24, Submitted



- NLR properties mapped in a large sample of  $\sim 330$  luminous quasars from SDSS DR16 including 73 BALs and 120 mini-BALs
- Broad line region (CIV) and narrow-line region ([OIII]) outflow velocities are correlated (albeit with significant scatter). Correlation not purely driven by luminosity, which suggests BLR winds are able to propagate and influence properties on up to kpc scales
- No difference in the NLR properties between BALs and non-BALs

# MOLECULAR GAS IN LUMINOUS Z=2 QUASARS

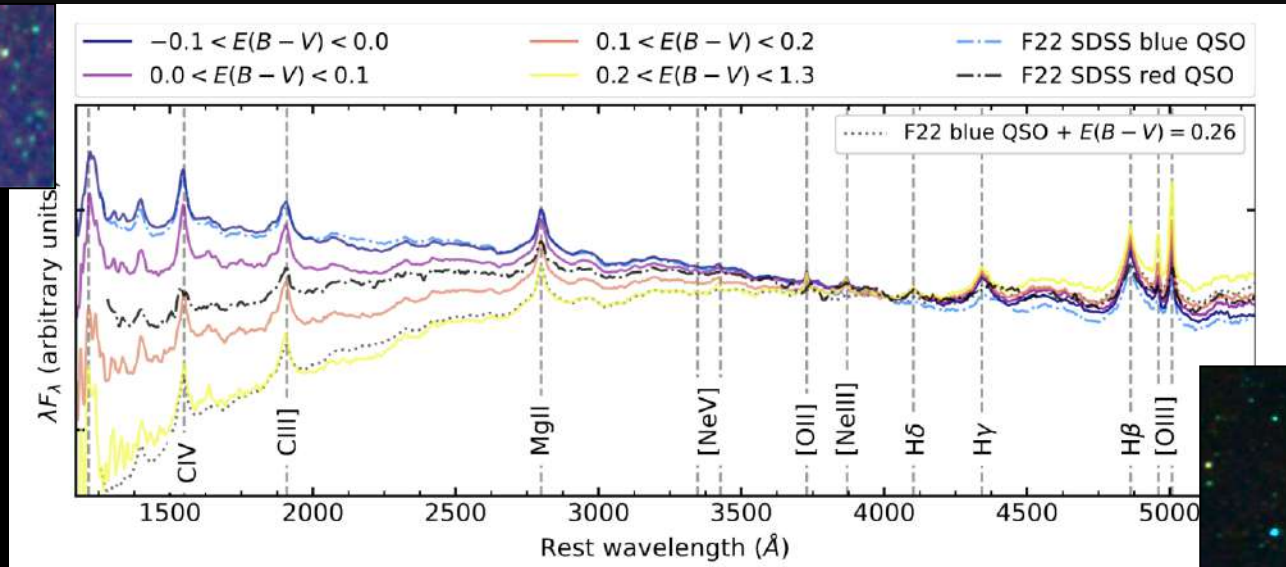
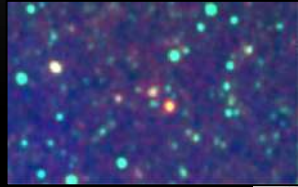


- ALMA Large Program to observe CO(3-2) in quasars across the CIV space – around 30 new quasars observed + archival data = 40 quasars at z=2
- No strong correlation in gas properties and UV line properties. Maybe tentative evidence that high ionizing SEDs are associated with gas-poor host galaxies while strong BLR outflows associated with gas-rich hosts.
- Detailed analysis ongoing

Molyneux, MB, Temple+ in prep

# DUSTY AND RED QUASARS

Fawcett+22

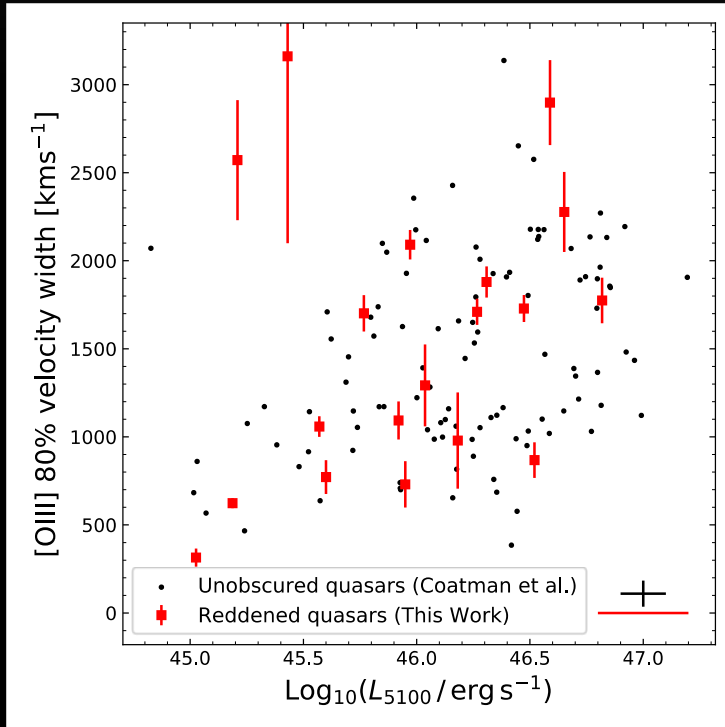


Sizeable populations of red/dusty broad-line (Type 1) quasars discovered using ground-based optical and infrared imaging and spectroscopic surveys (e.g. MB+12,13,15, Glikman+12,22; Fawcett+22, 23; Assef+15, LaMassa+17, Hamann+17, Temple+19...)

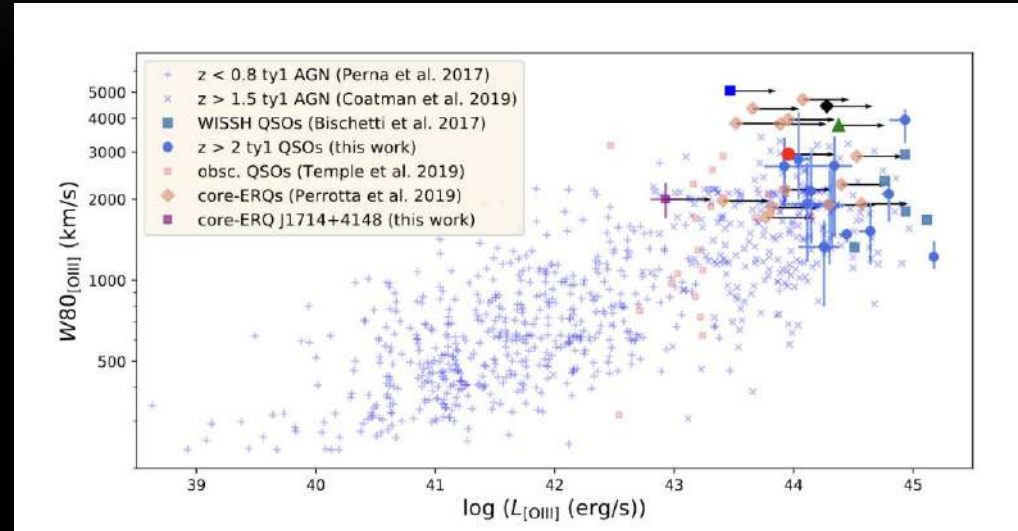
JWST now uncovering a very significant population of red AGN in the early Universe (e.g. Labbe+23, Greene+23) – 33% AGN among  $z_{\text{phot}} > 5$  selected galaxies!

# [OIII] OUTFLOWS IN DUSTY QUASARS

Temple, MB+19



Villar-Martin+18

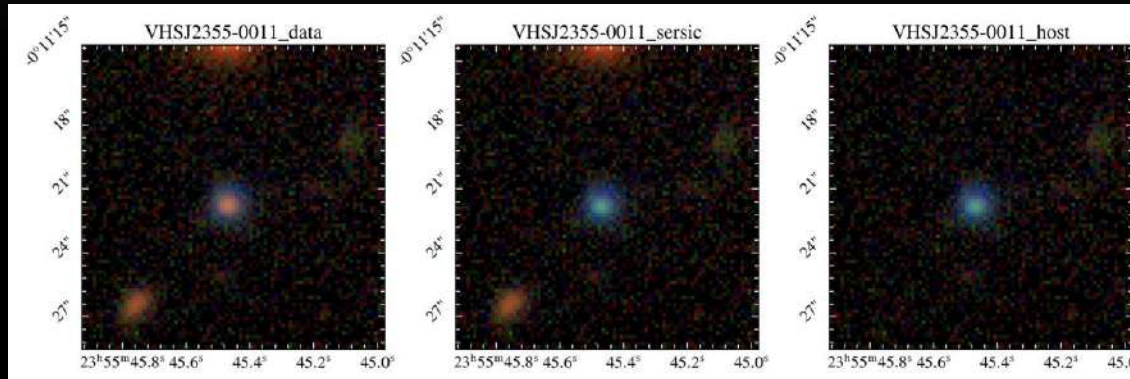


No evidence to suggest red/dusty quasars have stronger outflows compared to luminosity-matched blue quasars at similar redshifts



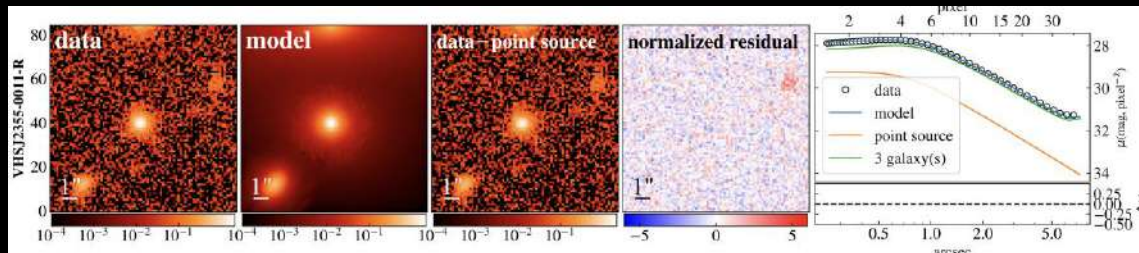
# HOST GALAXIES OF RED QUASARS WITH HSC

Tang, MB+ in prep



HSC imaging ( $\sim 0.5''$  seeing) now allows us to resolve host galaxy emission in  $z=2$  red quasars

>50 red quasars where we can already see resolved host galaxy emission in current imaging surveys (HSC, DECam)



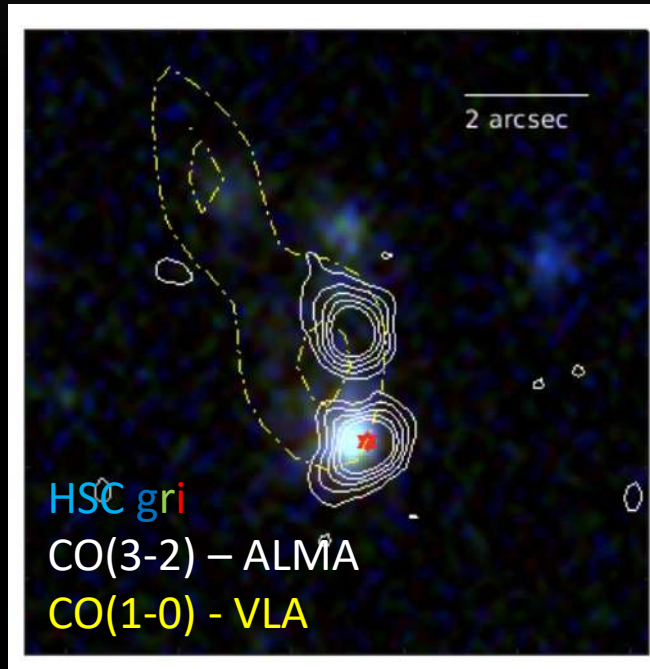
In some cases we even see evidence for mergers

See also [Urrutia+08](#), [Glikman+15](#), [Zakamska+19](#) for results from space-based imaging

LSST+Euclid will provide exquisite imaging for many high- $z$  quasar systems allowing host galaxies to be studied in detail in statistical samples

# HOST GALAXIES OF RED QUASARS FROM ALMA

MB+17, 18, 21



Detailed observations of a red ( $A_V \sim 5$  mag) quasar at  $z=2.566$

Multiple CO/CI transitions detected with ALMA and VLA

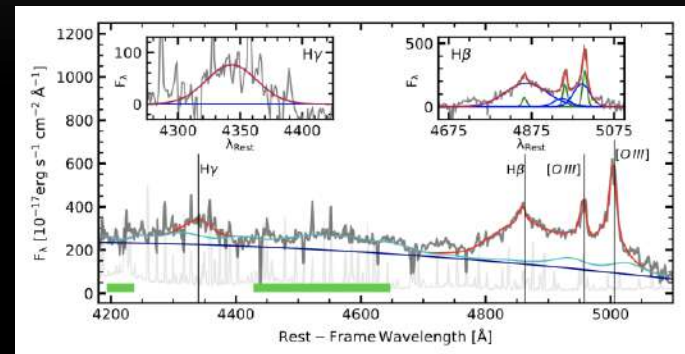
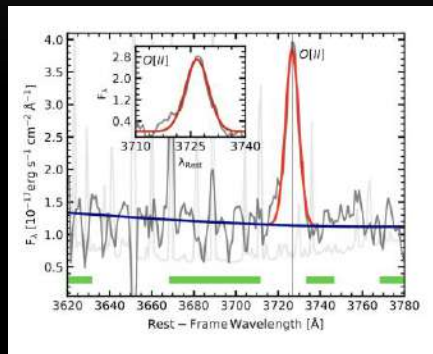
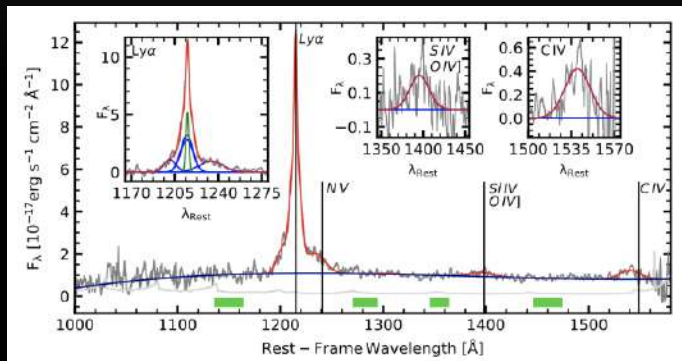
Major merger – 1:3 mass ratio, physical separation of  $\sim 15$  kpc, velocity separation  $\sim 170$  km/s

Extended rest-UV emission seen in HSC



# XSHOOTER VIEW OF A DUSTY RED QUASAR

Stepney, MB, Tang+ in prep



10 hr Xshooter spectrum of a dusty red quasar, ULASJ2315 at  $z=2.56$  covering rest-frame UV to rest-frame optical

- Broad Balmer emission lines -> massive black hole
- Broad blueshifted CIV emission -> strong outflows in quasar broad-line region
- Blue wings in [OIII] emission -> strong outflows affecting the narrow-line region gas
- [OII] emission – star formation in host galaxy
- Faint UV continuum inconsistent with dust-reddened quasar – scattered/leaked quasar light + star formation?
- Very strong, high EW Ly $\alpha$  emission – star formation?

MOONS AGN Survey (GTO) will obtain spectra for thousands of dusty/red quasars starting 2024/25

# SUMMARY

- Current and future wide-field imaging and spectroscopic surveys provide rich datasets within which to explore correlations between accretion and multi-phase, multi-tracer outflow properties in luminous, high- $z$  quasars
  - Dusty and red quasars offer a unique opportunity to simultaneously study black-hole and host galaxy properties in luminous quasars at cosmic noon
  - Current and forthcoming surveys and facilities e.g. DESI, WEAVE, 4MOST, VLT-MOONS, JWST, ALMA etc. will allow lots of progress in this area
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